

Imaging granular matter: some observations and challenges



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Osvanny Ramos

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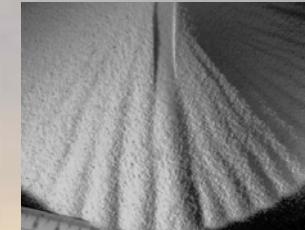
Øistein Johnsen

Knut J. Maløy

Hans Herrmann

Sept-Oct. 2004

Revolving rivers



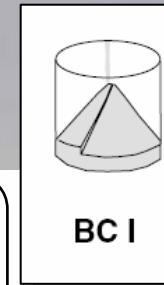
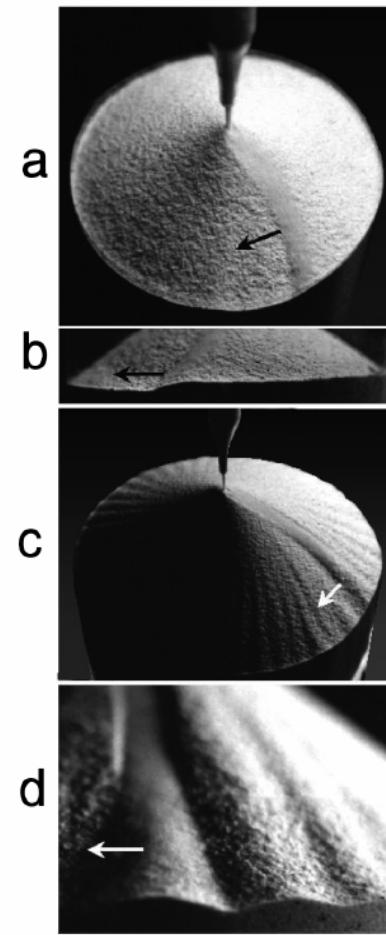
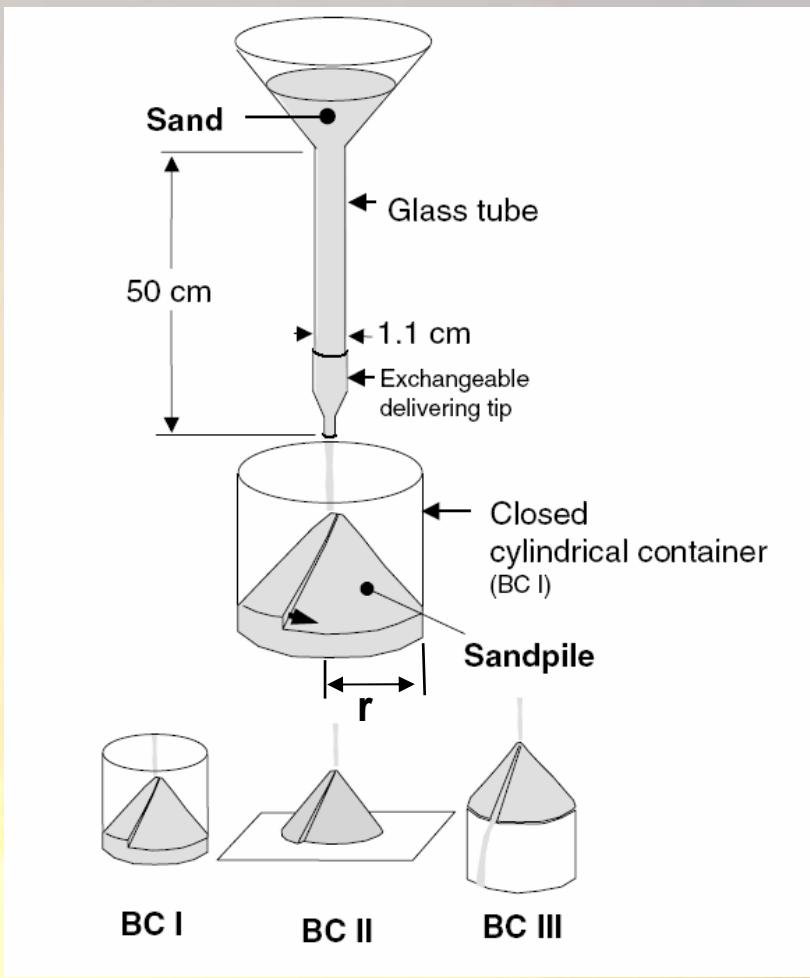
Uphill bumps



Other



Revolving rivers: introduction

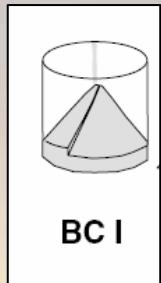


BC I

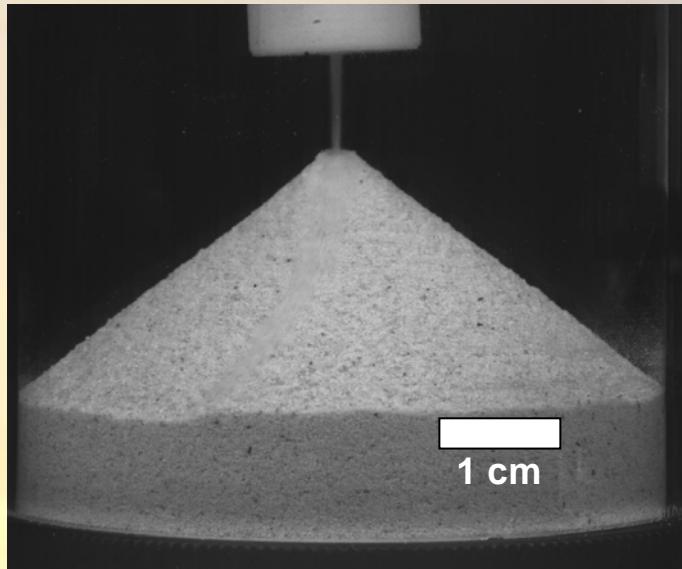
**Continuous
rivers**
(small radius)

**Intermittent
rivers**
(bigger radius)

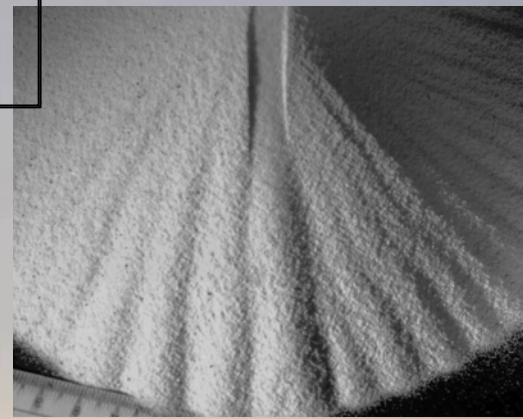
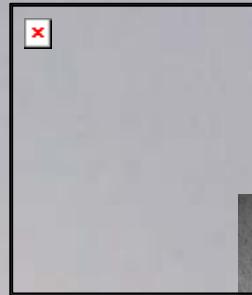
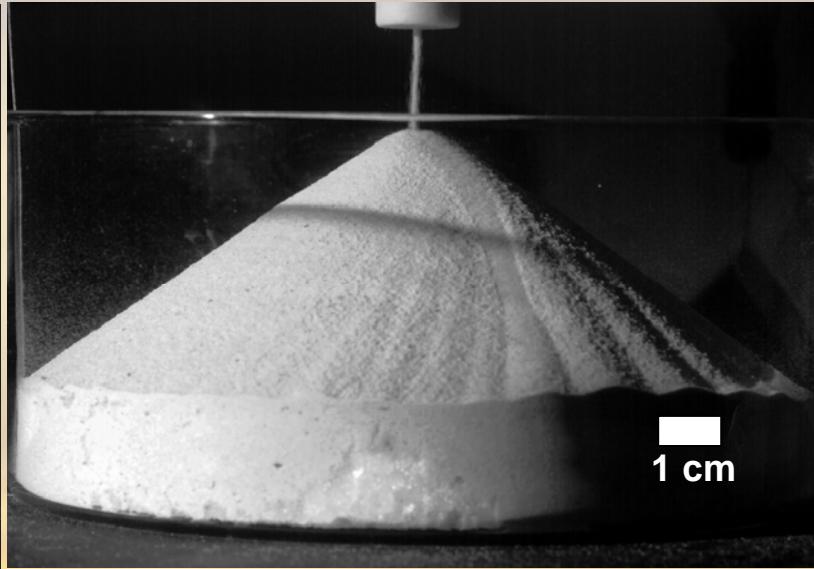
Revolving rivers: introduction



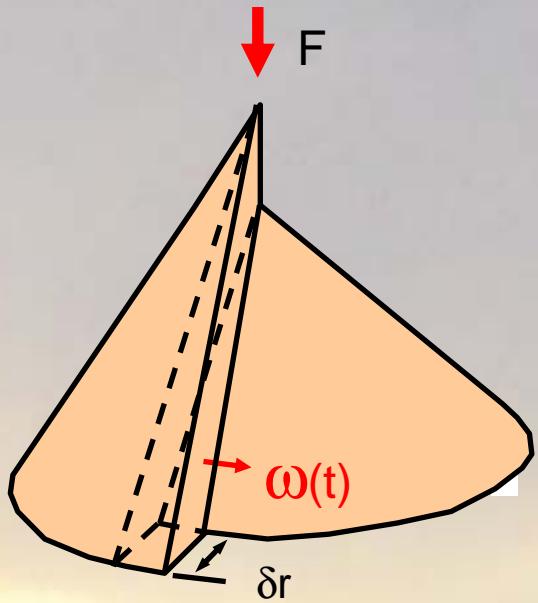
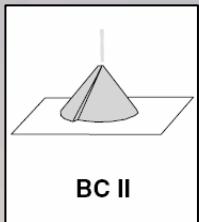
Continuous regime



Intermittent regime



Revolving rivers: a simple, old model



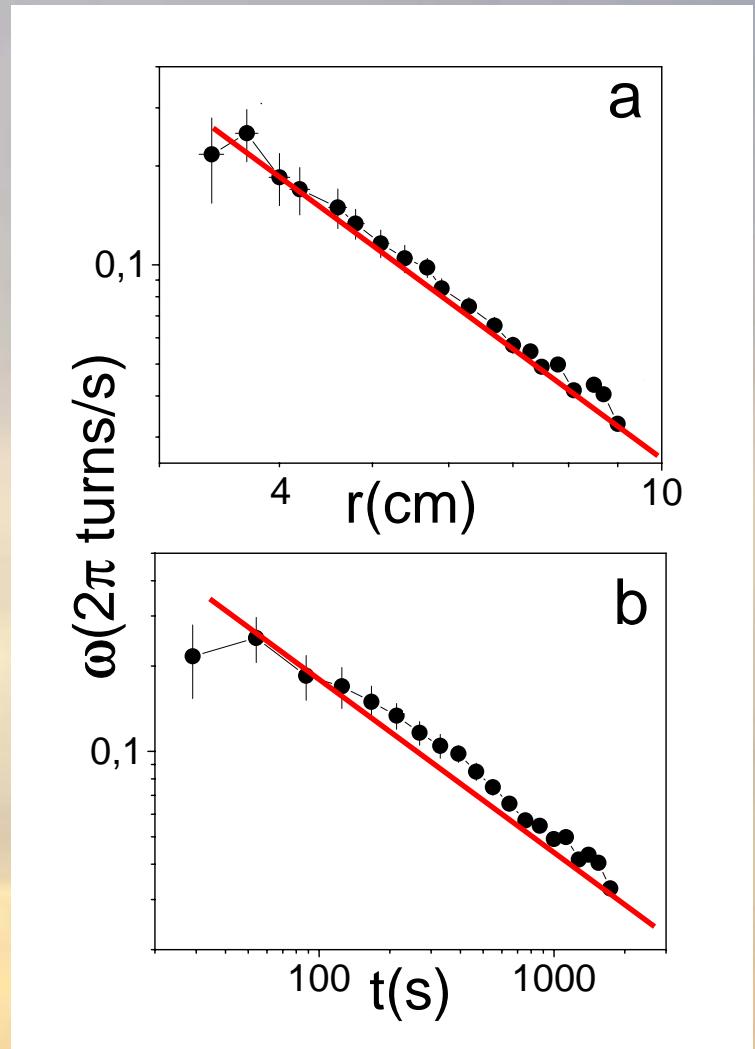
$$\omega = \frac{2F}{\tan \theta} \frac{1}{\delta r} \frac{1}{r^2}$$

$$\omega = \frac{2}{(\tan \theta)^{1/3}} \left(\frac{\pi}{3} \right)^{2/3} F^{1/3} \frac{1}{\delta r} \frac{1}{t^{2/3}}$$

$$F = 0,4 \text{ cm}^3 / \text{s}$$

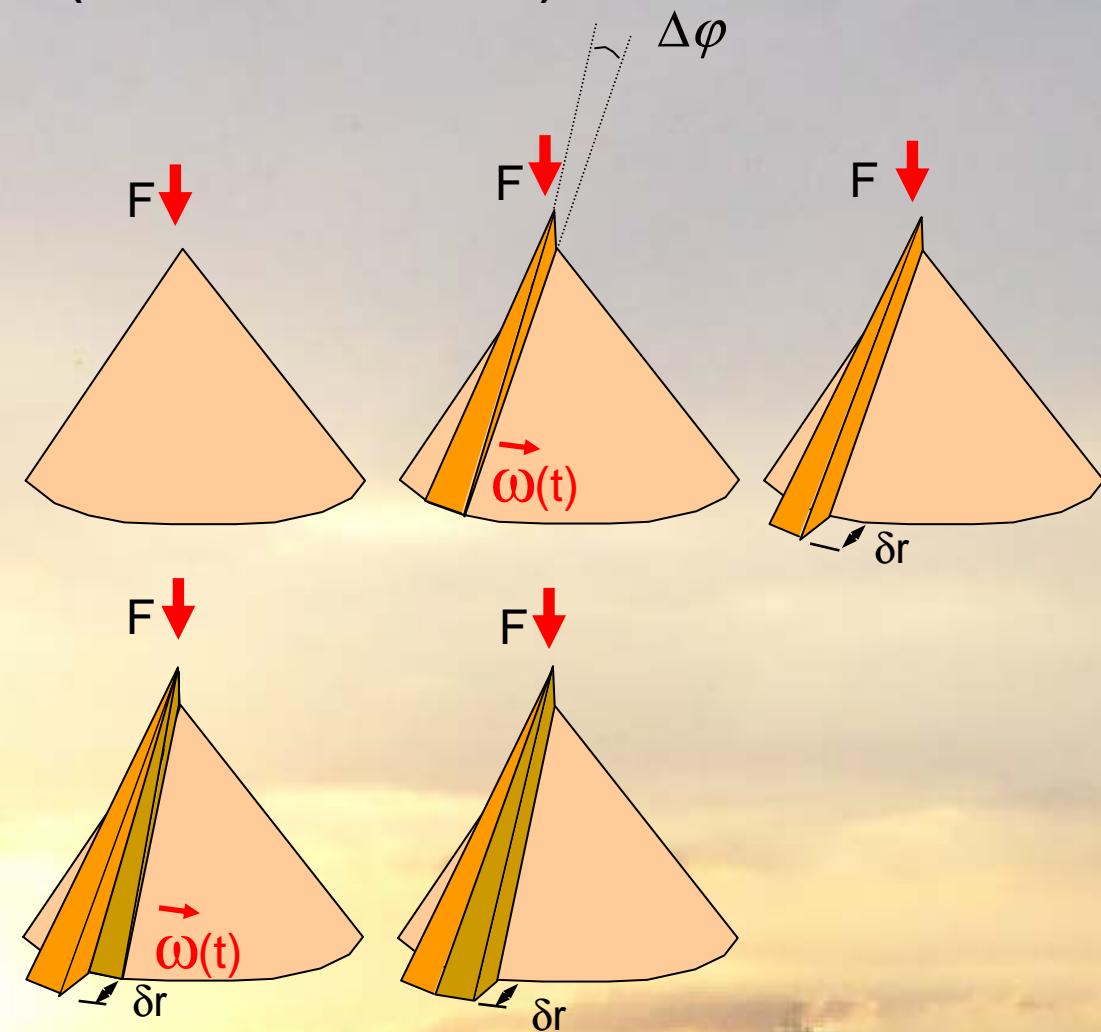
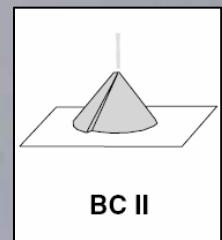
$$\theta = 33^\circ$$

$$\delta r = 0,4 \text{ cm}$$



Revolving rivers: the tickling model

(Herrmann & Toussaint)



$$\omega = 76 \frac{F \cos^2 \theta}{\tan\left(\frac{3}{2} \Delta\varphi\right)} \frac{1}{r^3}$$

Assumption No. 1

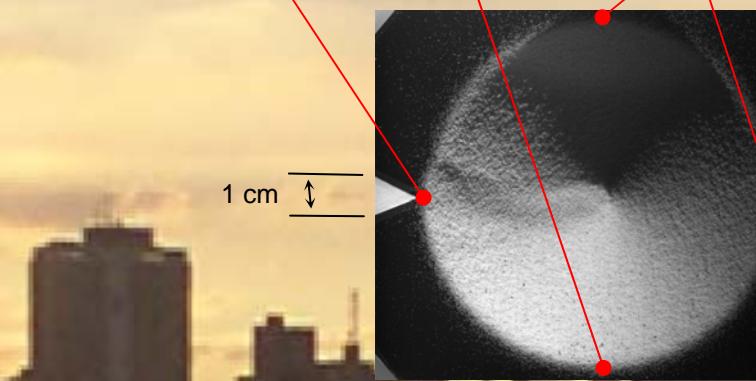
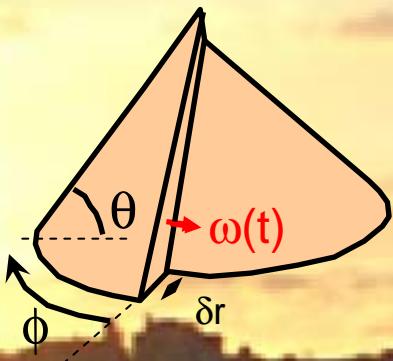
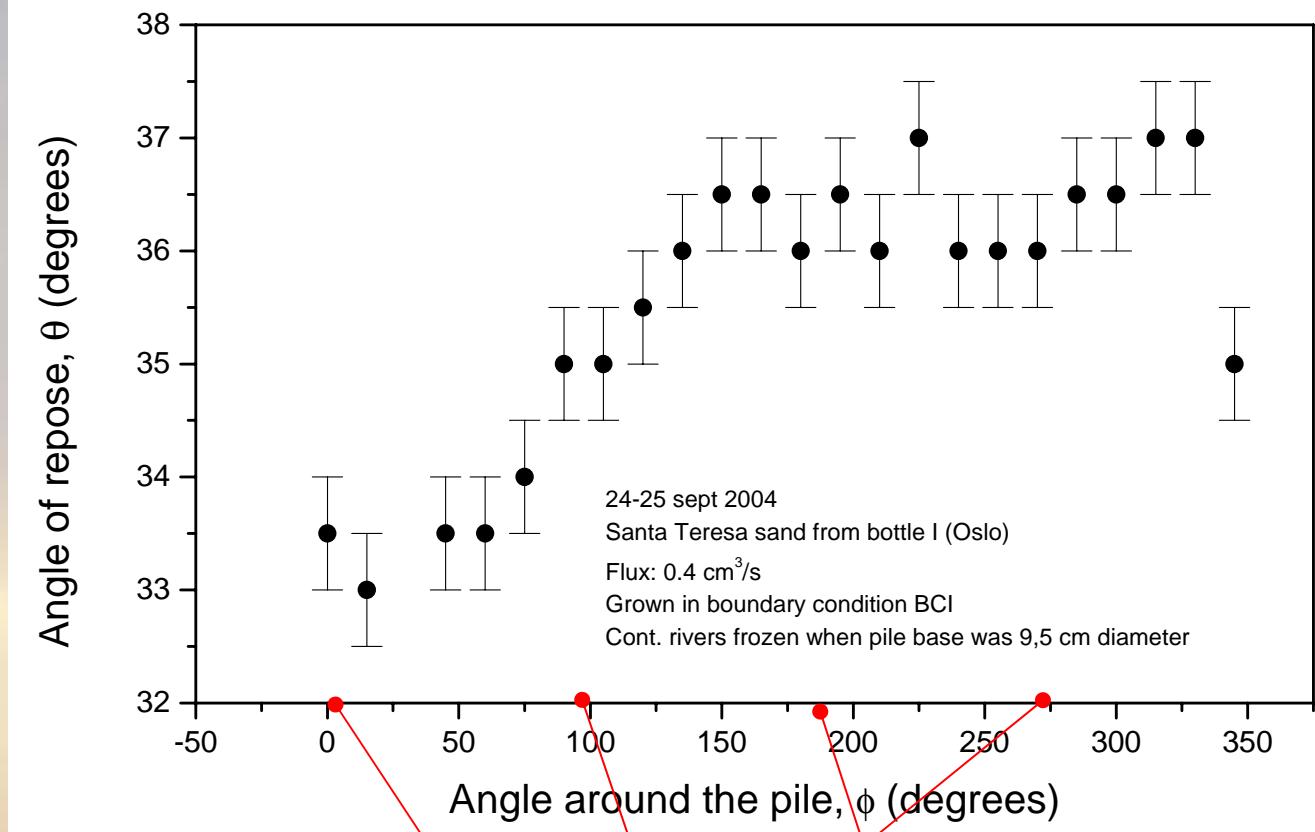
$$\Delta\varphi = \text{const.} \Rightarrow \omega \propto \frac{1}{r^3} \text{ (wrong!)}$$

Assumption No. 2

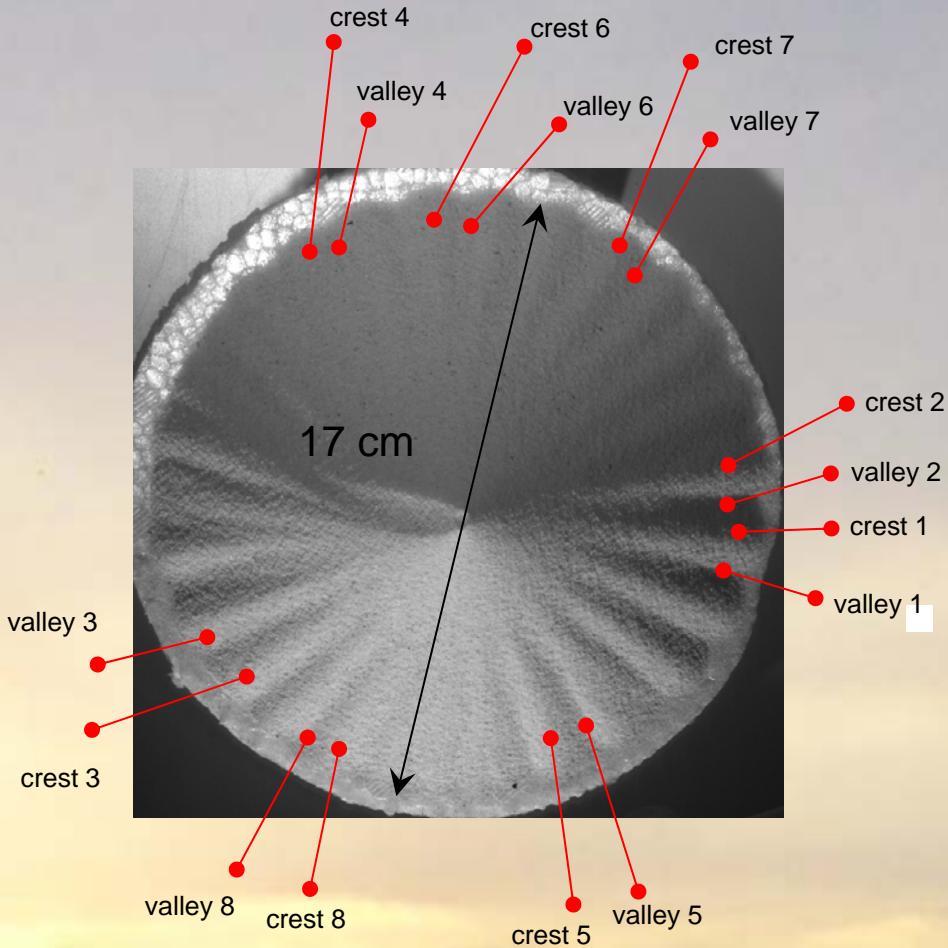
$$\delta r = \text{const.} \Rightarrow \Delta\varphi \propto \frac{1}{r};$$

$$\omega = \frac{2F}{\tan \theta} \frac{1}{\delta r} \frac{1}{r^2} \text{ (old, correct formula)}$$

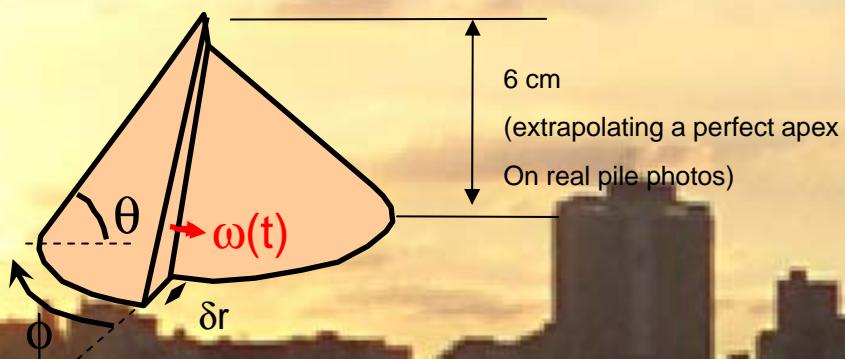
Continuous rivers: angular realities



Intermittent rivers: angular realities

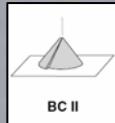
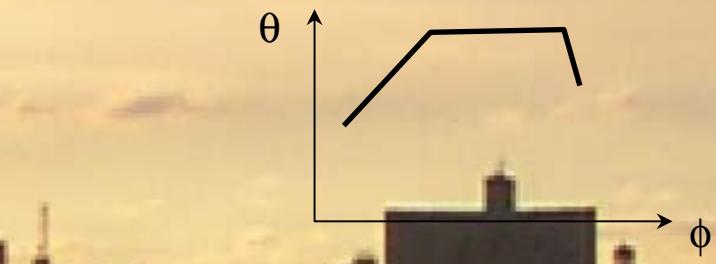
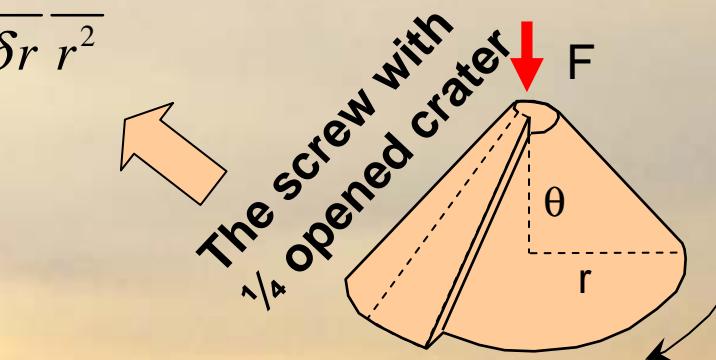
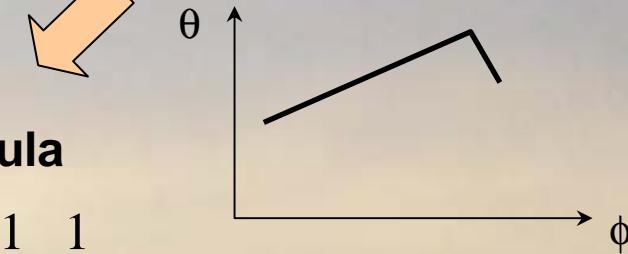
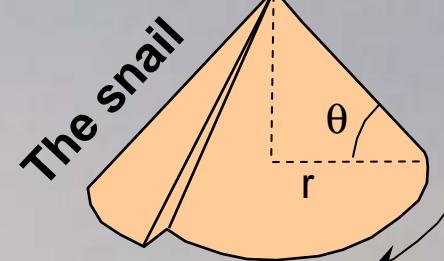
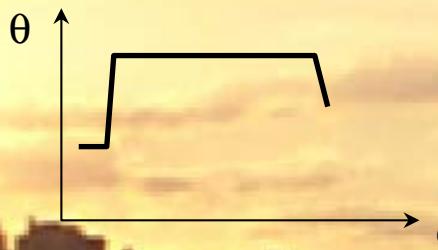
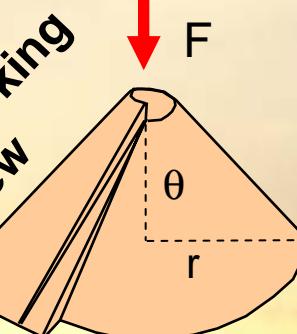
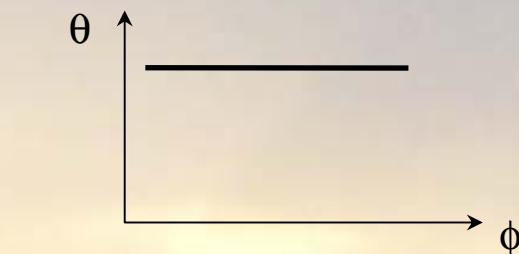
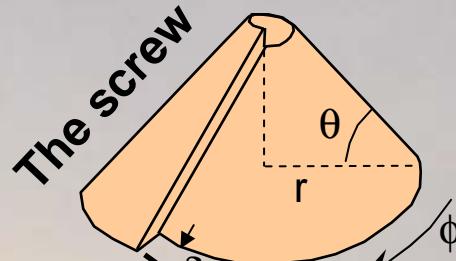


#	θ_{valley}	$\langle \theta_{\text{valley}} \rangle$	θ_{crest}	$\langle \theta_{\text{crest}} \rangle$
1	37		35	
2	37		36	
3	38		36	
4	37	37,1	36	36,1
5	37,5	$\pm 0,16$	36	$\pm 0,20$
6	37		36,5	
7	37		36	
8	36,5		37	



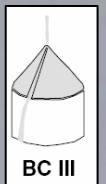
Revolving rivers: a zoology of models

(Toussaint)



Old formula

$$\omega = \frac{2F}{\tan \theta} \frac{1}{\delta r} \frac{1}{r^2}$$

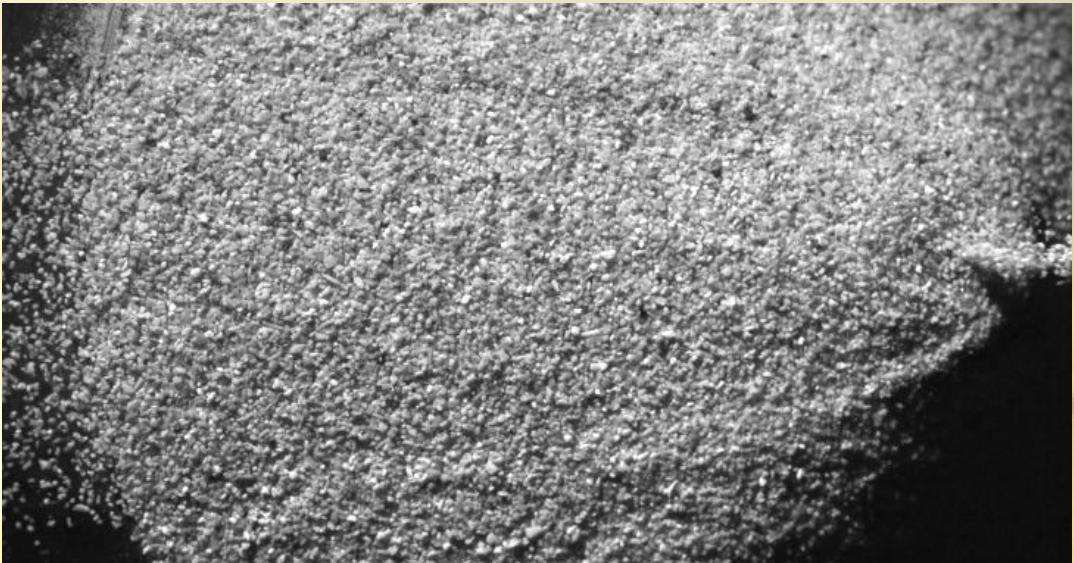


Why rivers, in the first place?

-- Spontaneous vs. Non-spontaneous symmetry breaking



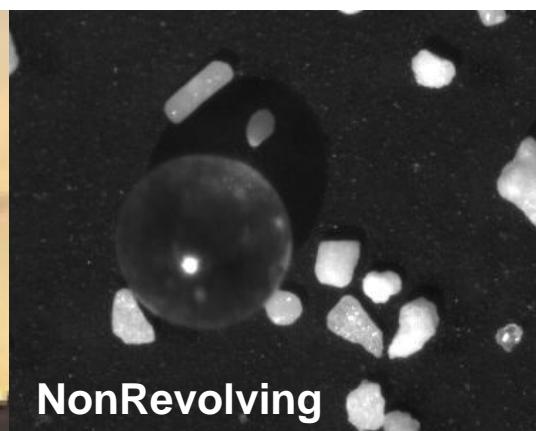
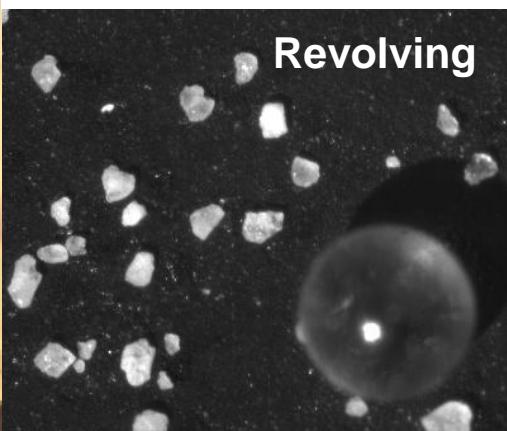
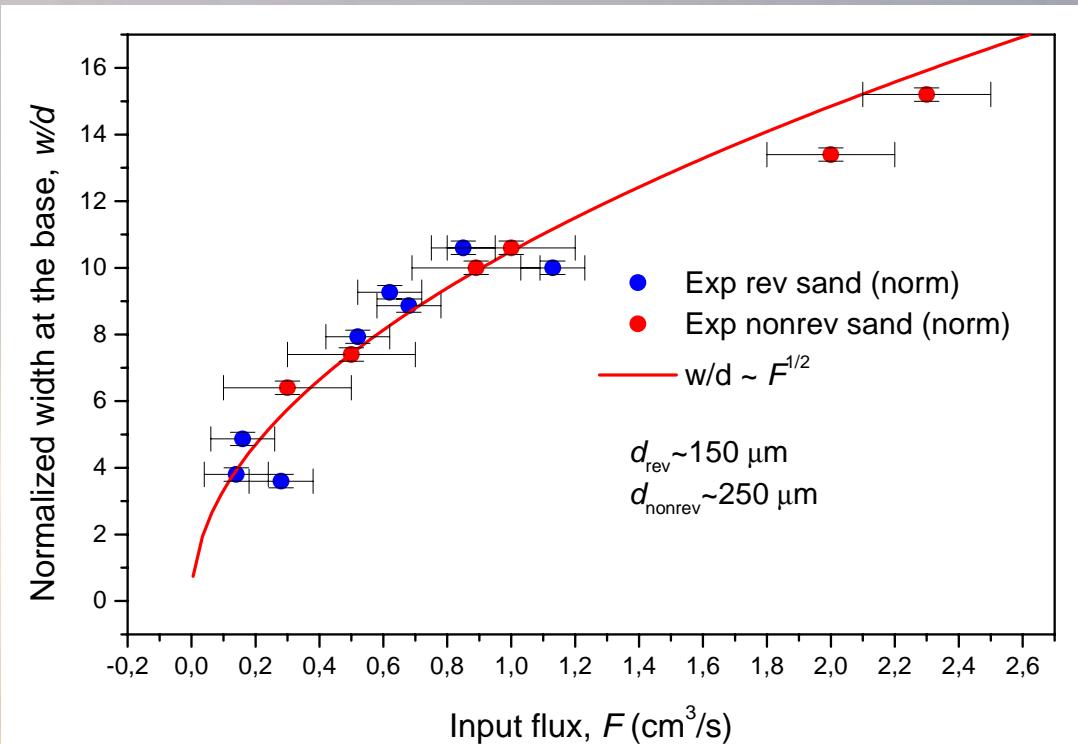
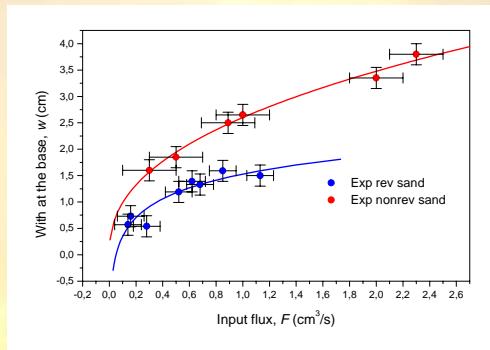
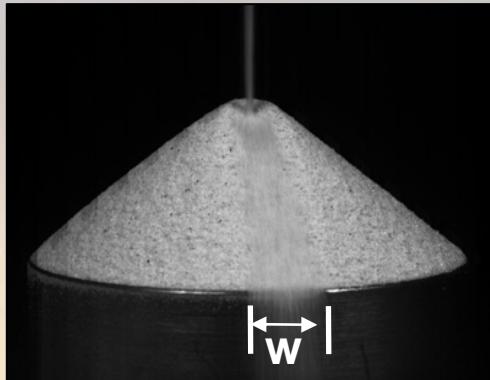
Revolving sand
(Santa Teresa)



Non-revolving sand
(Varadero)

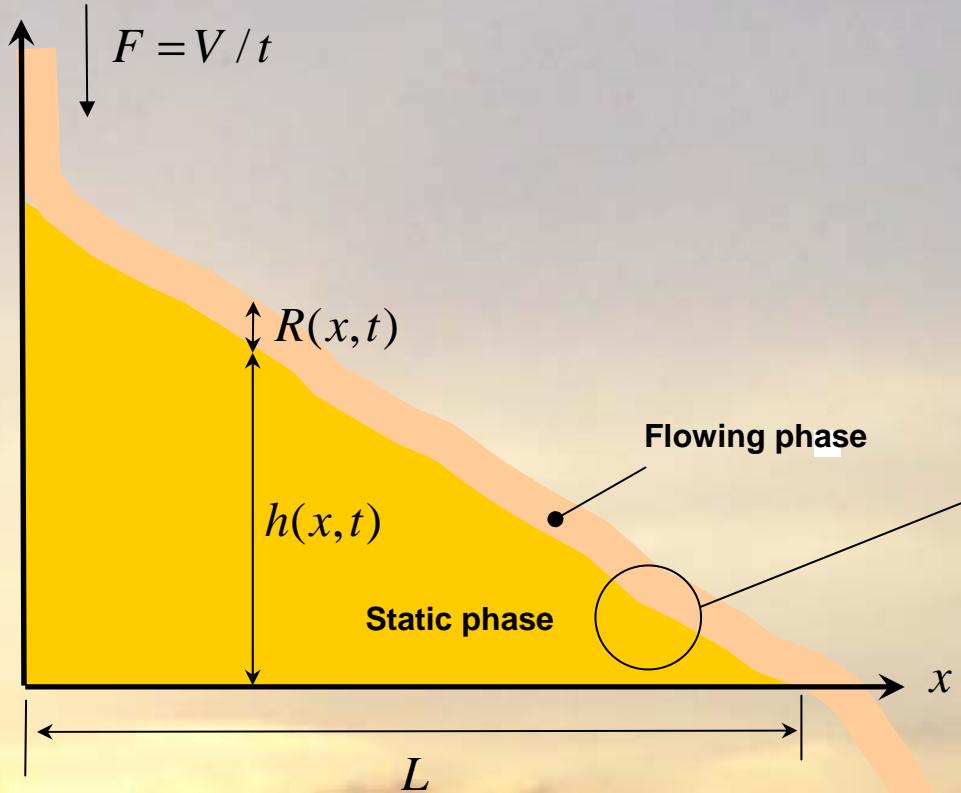
Why rivers, in the first place?

-- The geometry of the flow: a possible hint



Flowing sand: the BCRE model

Bouchaud, Cates, Prakash & Edwards, J. Phys. France 4 (1994) 1383



$$\frac{\partial R}{\partial t} \approx -\frac{\partial h}{\partial t} + v \frac{\partial R}{\partial x}$$

$$\frac{\partial h}{\partial t} \approx \gamma R(\theta_n - \theta)$$

$$v_{uphill} = \gamma R = \sqrt{\frac{g}{d}} R$$

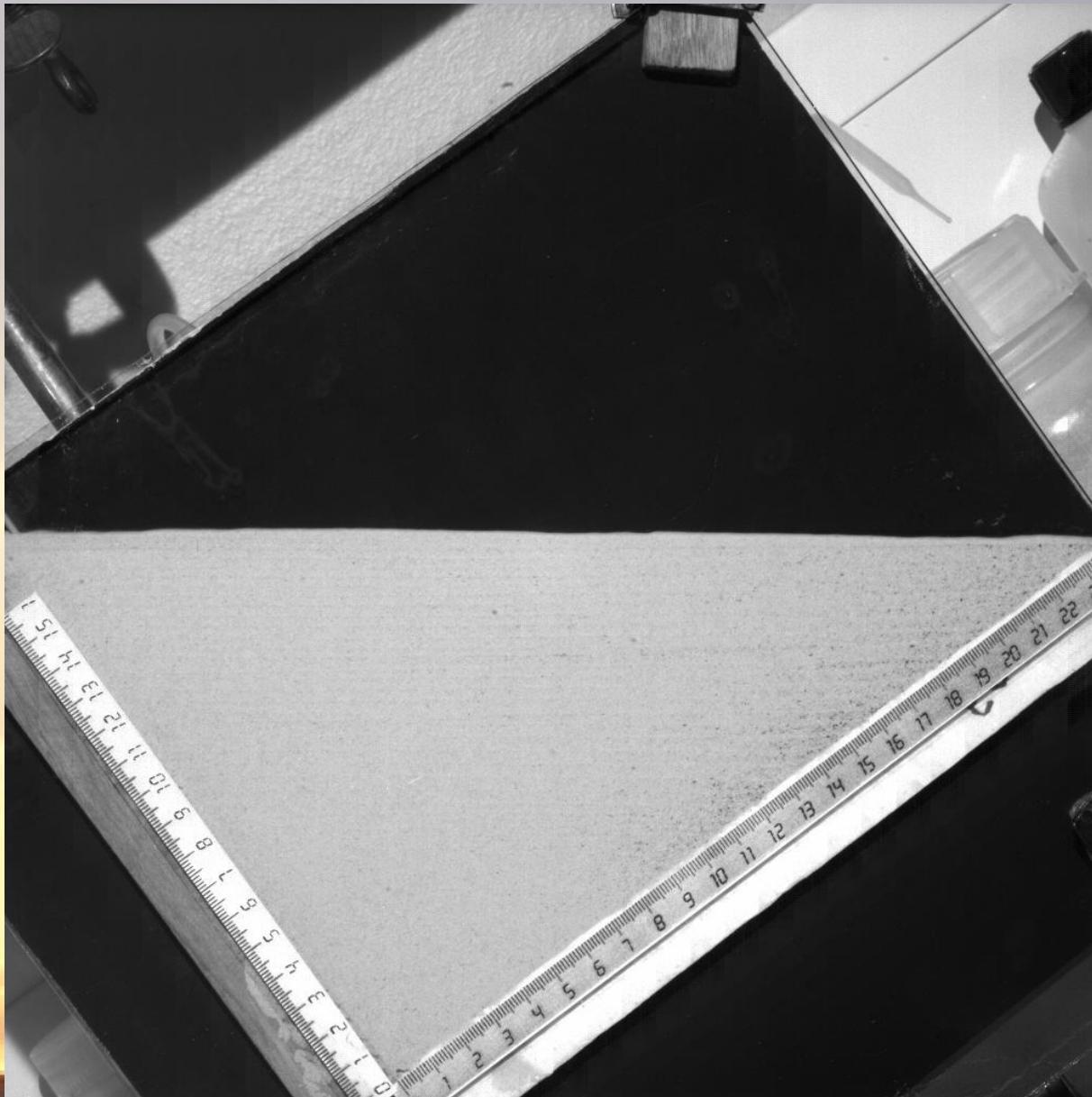
For spheroidal grains and
"average levels
of inelastic collisions"

Accident in slope ("bump")

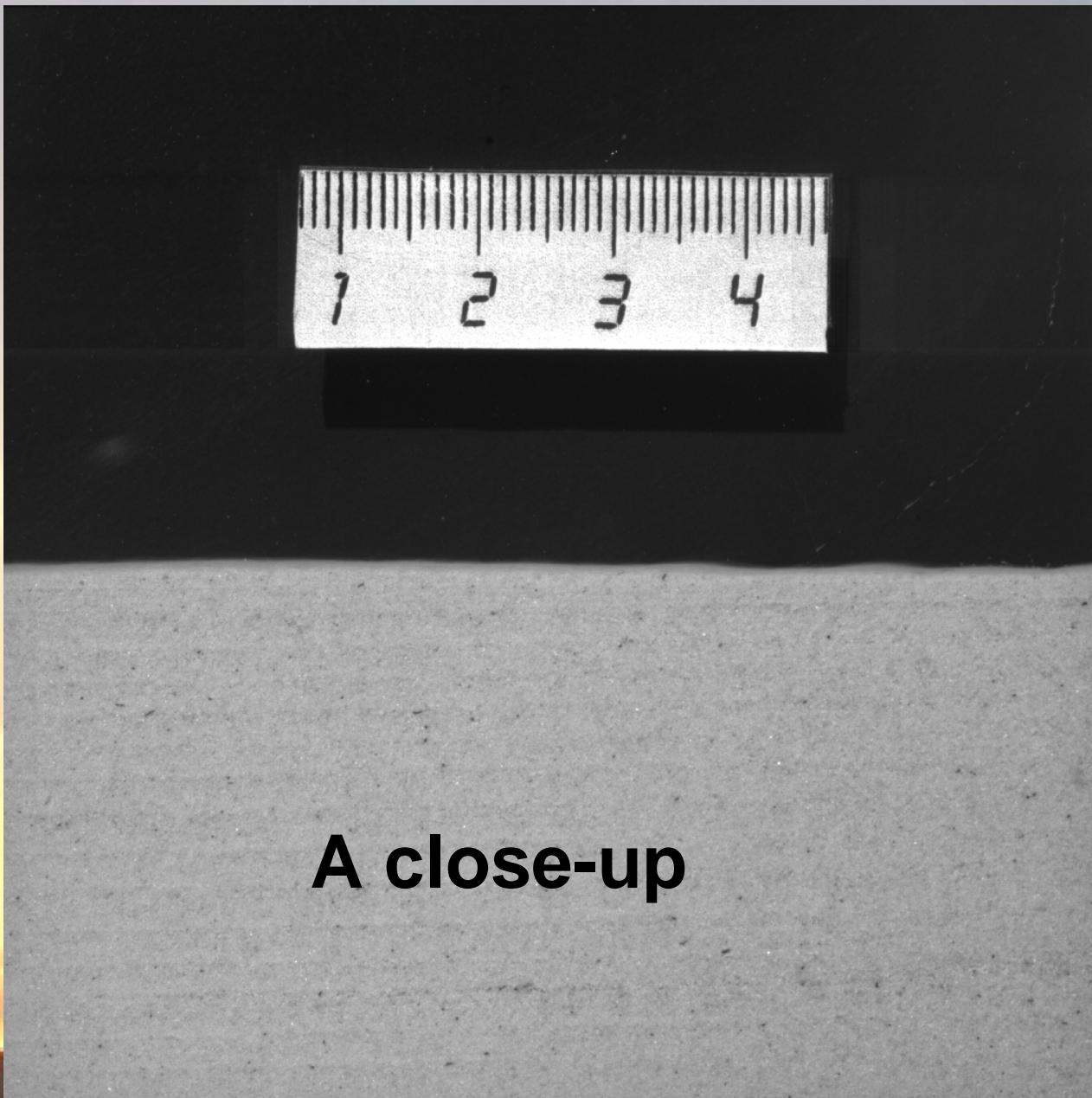
Following the simplification by Boutreux, Raphael and DeGennes, Phys. Rev. E 58 (1998) 4892

$$\frac{\partial h}{\partial t} \approx v_{uphill}(\theta_n - \theta), \quad v_{uphill} \square \frac{Q}{L(\theta_n - \theta)}$$

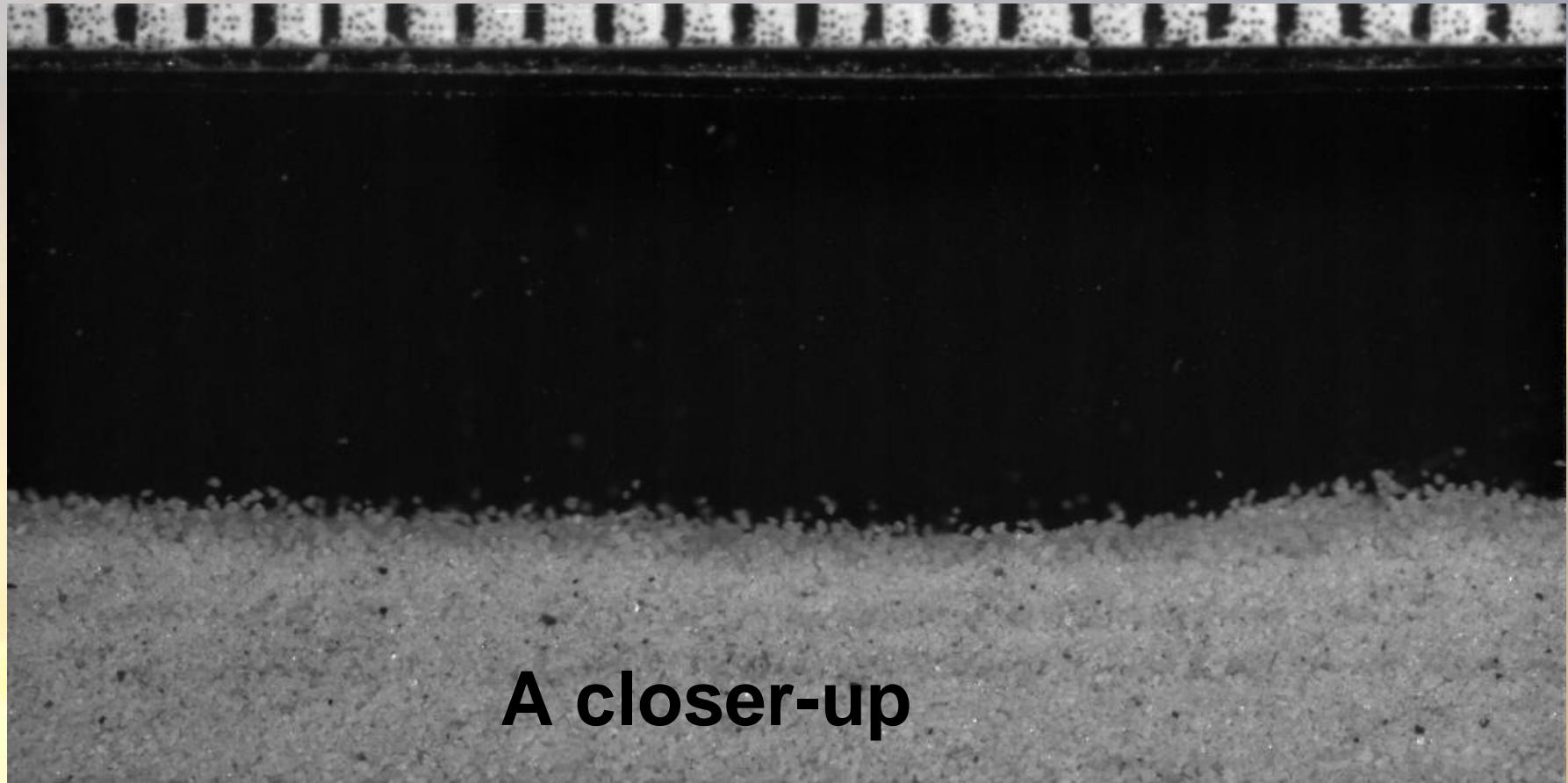
Experimental observation of "uphill" bumps



Experimental observation of "uphill" bumps



Experimental observation of "uphill" bumps



A closer-up

Checking some BCRE predictions

In our case,

$$L = 240\text{mm}$$

$$\theta_n - \theta \approx 3^\circ$$

$$d \approx 100\mu\text{m}$$

$$R \approx 1,5\text{mm}$$

So the BCRE theory gives

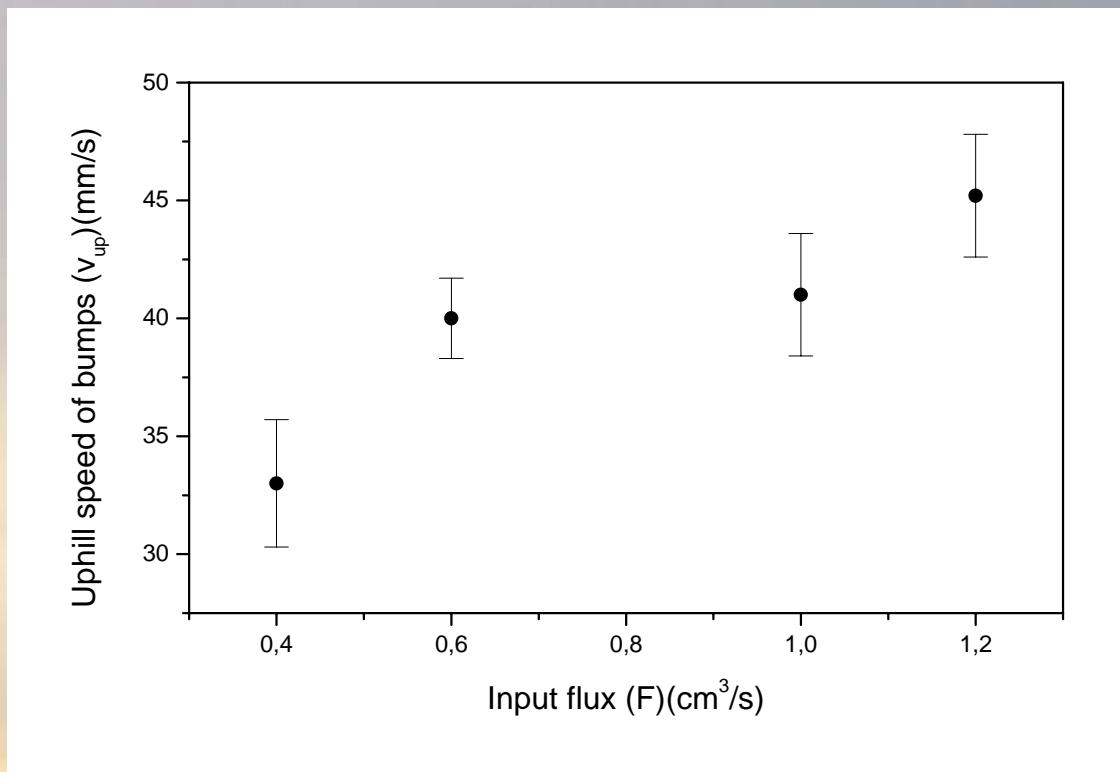
$$v_{uphill} = \sqrt{\frac{g}{d}} R \approx 30\text{cm/s} \dots \text{too big!}$$

$$\gamma \approx 32\text{s}^{-1}$$

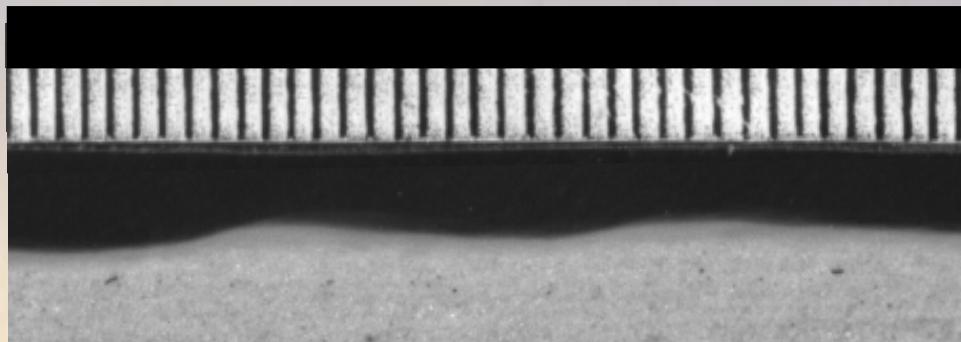
and the BRDG one gives

$$v_{uphill} \square \frac{F_{2D}}{L(\theta_n - \theta)} \approx \frac{F_{3D}}{Lw(\theta_n - \theta)} \approx 2F_{3D}$$

Our measurements



Moving bumps in heaps and tubes: a misleading resemblance?

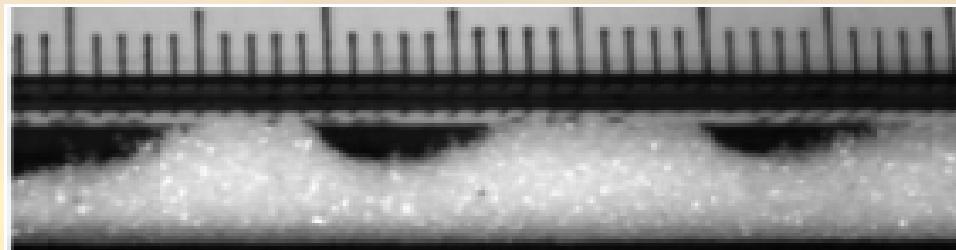


$$d \approx 100\mu m$$

$$R \approx 1,5 mm$$

$$\theta \approx 33^\circ$$

$$v_{uphill} \square 20 - 50 mm/s$$



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PHYSICAL REVIEW LETTERS

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Structure Formation and Instability in a Tube of Sand

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(Received 1 March 2001; revised manuscript received 2 August 2001; published 10 September 2001)

$$d \approx 180\mu m$$

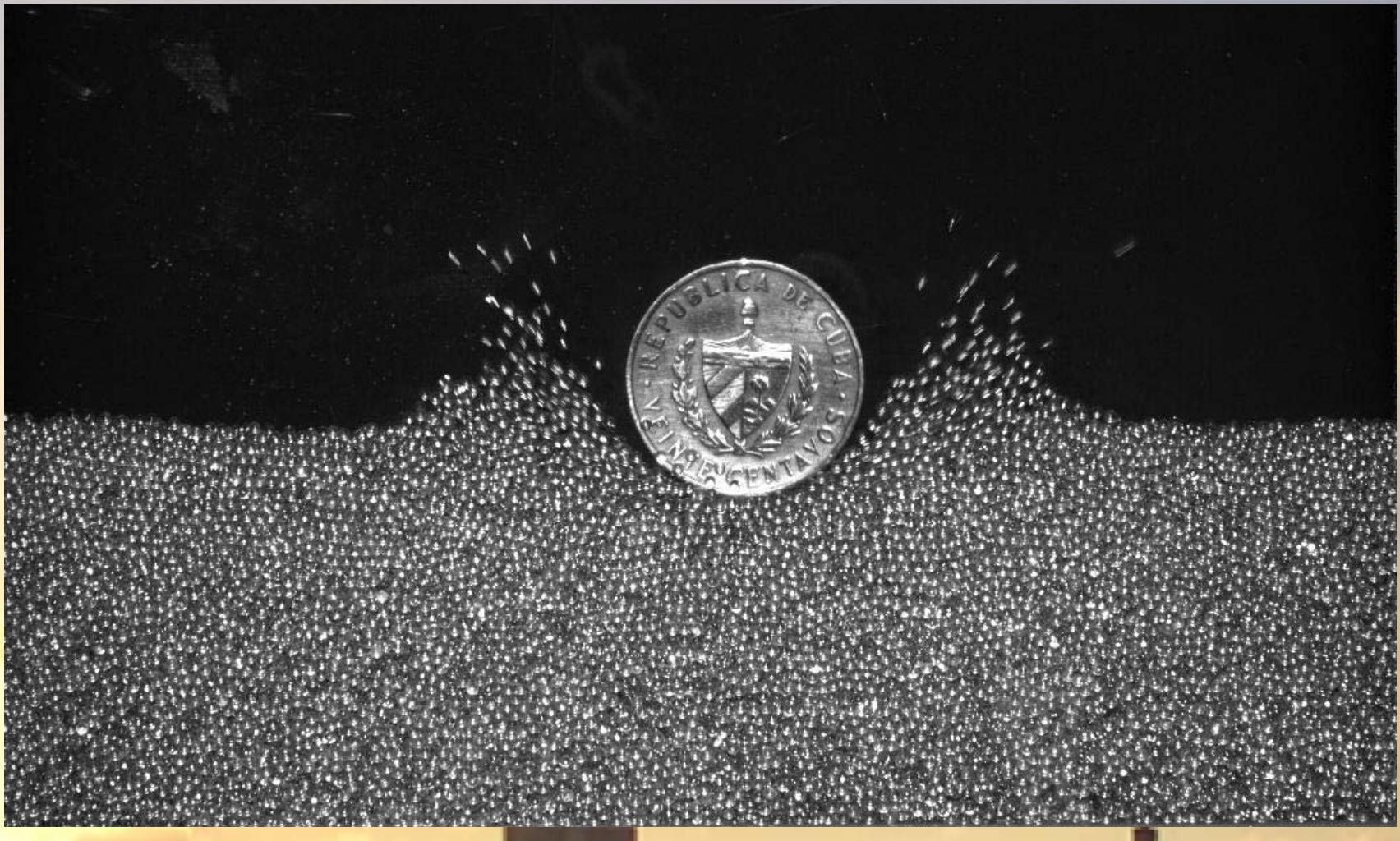
$$R \approx 1,5 mm$$

$$\theta \approx 70^\circ$$

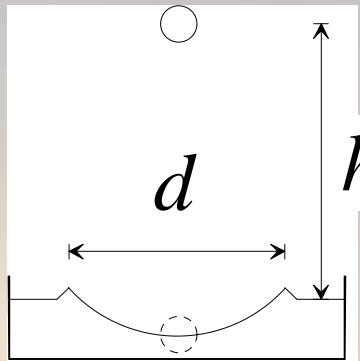
$$v_{uphill}^{bubbles} \square 100 - 500 mm/s$$

Other stuff:

Introducing a really inexpensive experiment



Impact cratering: a simple exercise



Impact-explosion analogy (3D)

$$d \propto E^{1/3}$$

Gravitational energy (3D)

$$d \propto E^{1/4}$$

Gravitational energy (Ideal 2D)

$$d \propto E^{1/3}$$

